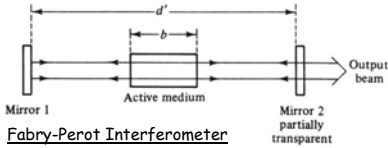


Optical Amplification

- We can enhance the *intensity* of the emission by using an *optically resonant cavity*:

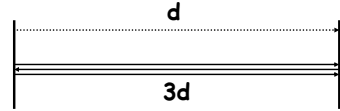


Fabry-Perot Interferometer

Get constructive interference if: $d' = n\lambda/2$

Constructive Interference

- Waves are *perfectly in-phase* if the difference in the distance they travel is an **integer multiple of the wavelength**:



$$3d - d = 2d = n\lambda$$

$$d = n\lambda/2$$

EMR from a Laser is Unusual

- **Spatial Coherence**
 - all photons are *in-phase*
 - can attain very high *power density*
 - very low *beam divergence*
- **Spectral Coherence**
 - highly monochromatic
- **High-Intensity**
 - very high power over very small $\Delta\lambda$

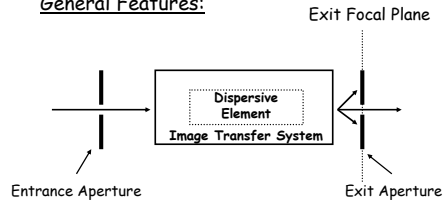
Very rarely used for *absorption spectrophotometry*, but important sources for *fluorescence and Raman spectroscopies*

Wavelength Selectors

- **Dispersive Devices**

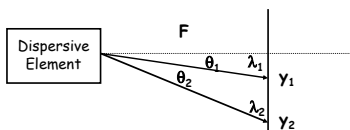
-separates EMR into individual λ -components

General Features:



Dispersion (Terminology)

- How do we quantify the spatial separation of wavelengths on the exit focal plane?



Angular Dispersion: $D_a = d\theta/d\lambda$ (property of dispersive element)

Linear Dispersion: $D = dy/d\lambda$ (property of dispersive device)

More Dispersive Terminology

- If $d\theta$ is small, it can be shown that:

$$D = F \times D_a$$

• More commonly, we will use:

Reciprocal Linear Dispersion (D^{-1}) = $1/D$

-example: typically around $0.1 - 20 \text{ \AA/mm}$ in UV/Vis

Effective Bandwidth:

$$\Delta\lambda_{\text{eff}} = D^{-1} \times w$$

$D^{-1} = 16 \text{ \AA/mm}$
 $w = 100 \text{ \mu m}$

$\Delta\lambda_{\text{eff}} = 1.6 \text{ \AA}$

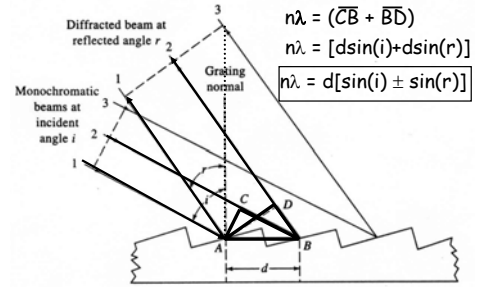
slitwidth

Diffraction Gratings

- Typically, a series of closely spaced *facets* ruled onto a *reflecting surface*
 - Spacing of facets must be comparable to λ of EMR
 - Parallel EMR rays striking adjacent facets will travel different distances
 - Constructive interference* occurs if the difference in the distance travelled by the two rays is an integer multiple of λ s
 - Constructive interference* will be a function of the *angles* (incident and reflection) and the *wavelength*

13

Let's see how it works!



14

Spectral Order

$$n\lambda = d[\sin(i) \pm \sin(r)]$$

spectral order

For fixed values of i and r , $n\lambda$ is constant.

Example: $n\lambda = 6000 \text{ \AA}$

n :	1	2	3	4
λ :	6000	3000	2000	1500

Special Case:

$n\lambda = 0$ (zeroth order)
ALL λ s reflected!

15

Dispersion for a Grating

- For a fixed angle of incidence (i):

$$n d \cos(r) = d(r)$$

So:

$$\frac{d(r)}{d\lambda} = \frac{n}{d \cos(r)}$$

D_o

If the angle r is kept small ($< 5^\circ$):

$$\frac{d(r)}{d\lambda} \approx \frac{n}{d}$$

wavelength independent

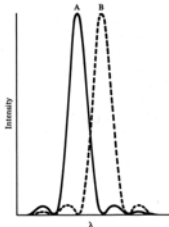
16

Spectral Resolution/Resolving Power

- We define *Resolution* (or *Resolving Power*):

$$R = \lambda_{\text{avg}} / \Delta\lambda$$

R is the resolving power needed to *just resolve* two spectral features:
Rayleigh Criterion



17

Resolving Power for a Grating

- So, in order to *just resolve* these two spectral lines:

$$\lambda_1 = 4501 \text{ \AA}$$

$$\lambda_2 = 4499 \text{ \AA}$$

We need an instrument with a *resolving power* of:

$$R = 4500 / 2 = \underline{\underline{2250}}$$

- The *resolving power* of a diffraction grating:

$$R = nN$$

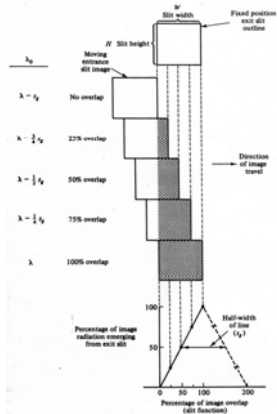
spectral order

facets illuminated on grating surface

18

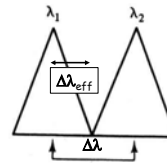
Slitwidth-Limited Resolution

- If $\Delta\lambda_{\text{eff}}$ is greater than the spectral linewidth, then resolution is controlled by the slitwidth
- Consider the overlap function of the entrance and exit slits of a scanning spectrometer:



Baseline-Resolution?

- Under slitwidth-limited resolution conditions, baseline resolution is attained when:



$$\Delta\lambda \geq 2\Delta\lambda_{\text{eff}}$$

So, $\Delta\lambda_{\text{eff}}$ must be smaller than the wavelength separation between the two spectral features (no more than $\frac{1}{2}\Delta\lambda$).

20

Effect of Slitwidth on Resolution

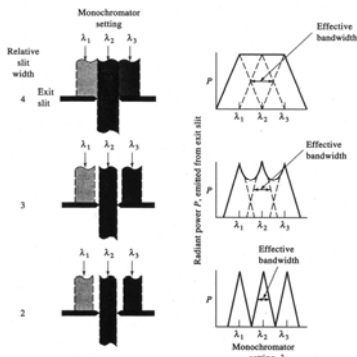
$$\Delta\lambda = \Delta\lambda_{\text{eff}}$$

"just resolved"

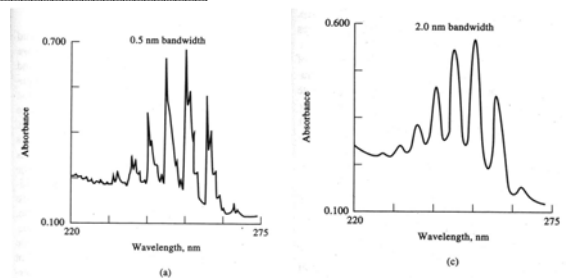
$$\Delta\lambda > \Delta\lambda_{\text{eff}}$$

$$\Delta\lambda = 2\Delta\lambda_{\text{eff}}$$

Baseline resolved



Effect of Slitwidth on Spectrum (Benzene)



22

How it all fits together

- Suppose we want to "just resolve" the following Iron doublet:

$$\lambda_1 = 3099.90 \text{ \AA} \quad R = 3099.935/0.07$$

$$\lambda_2 = 3099.97 \text{ \AA} \quad = \underline{44,000}$$

Suppose that we have a 100-mm wide grating ruled with 1200 gr/mm; it has a first-order resolving power of:

$$R = nN = (1)(100 \text{ mm})(1200 \text{ gr/mm})$$

$$R = \underline{120,000}$$

23

But, can we really resolve the two lines?

- Consider, now, the dispersion of the monochromator in which that grating is located:
 $D^{-1} = 16 \text{ \AA/mm}$

In order to just resolve the two lines:
 $\Delta\lambda_{\text{eff}} = 0.07 \text{ \AA}$

This requires a slitwidth of:

$$\Delta\lambda_{\text{eff}} = D^{-1} \times w$$

$$w = 0.07 \text{ \AA} / 16 \text{ \AA/mm} = 0.00438 \text{ mm}$$

$$w \approx \underline{4 \text{ \mu m}}$$

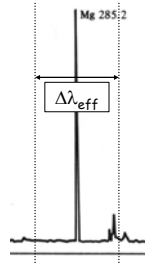
24

How Can We Improve Resolution?

- **Decrease Slitwidth (w)**
-limits light throughput
- **Operate in Higher Spectral Orders**
-limits light throughput (decreased efficiency)
 $D^{-1} \propto 1/n$
- **Increase Focal Length**
-limits light throughput (inverse square law)
 $D^{-1} \propto 1/F$

25

Resolution versus Light Throughput



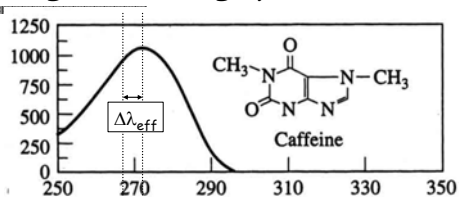
Case I:

$$\Delta\lambda_{\text{eff}} \gg \text{actual/linewidth}$$

- slitwidth limited resolution
- common with *atomic* spectra
- decreasing $\Delta\lambda_{\text{eff}}$ will **increase L/B** (!)

26

Resolution versus Light Throughput: II

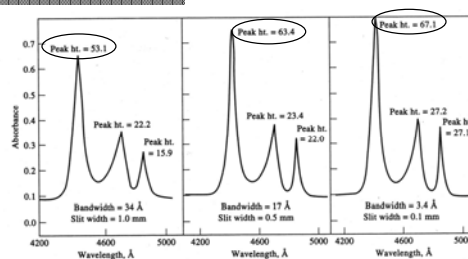


Case II: $\Delta\lambda_{\text{eff}} < \text{actual/bandwidth}$

- common with *molecular* spectra
- decreasing slitwidth, improves resolution, but *decreases* light throughput

27

Effect of Slitwidth on Absorbance

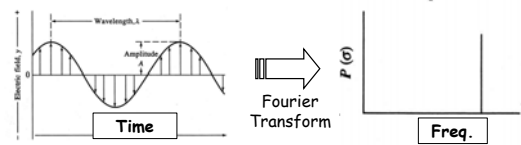


28

Non-Dispersive Methods

■ Fourier-Transform Interferometry

What if we could measure the *oscillating wavefunction* of EMR *directly*?



Time Domain

Frequency Domain

29

Direct Measurement: Feasible?

- Suppose we had EMR with $\lambda = 10 \mu\text{m}$

$$\text{Freq} = \nu = c/\lambda = \frac{3.00 \times 10^8 \text{ m/s}}{10 \times 10^{-6} \text{ m}}$$

$$\nu = 3 \times 10^{13} \text{ Hz}$$

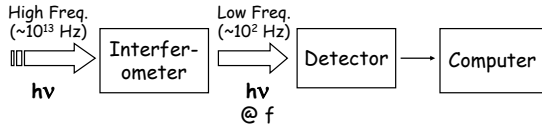
That's 1 cycle every 33×10^{-15} secs
(33 femtoseconds!)

Upshot: we *can't* measure the oscillating EMR field *directly* for *optical* radiation

30

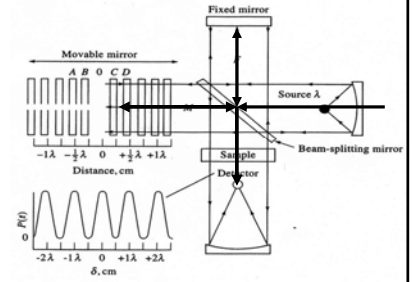
Enter Interferometry

- We need a signal that is *much* slower, so that it can be measured . . . How?



Michelson Interferometer

- EMR enters and hits *beamsplitter*
- Part goes to *fixed mirror*
- Part goes to *moveable mirror*
- Reflected beams recombine at *beamsplitter*

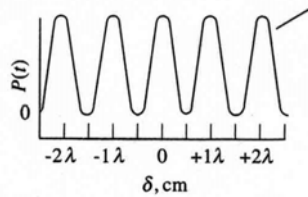


Resulting Interferogram

δ = pathlength difference (retardation)

$$\delta = 2(M-F)$$

$$\delta = 2x \text{ (mirror displacement)}$$



So, we get *maxima* when $\delta = n\lambda$ and *minima* when $\delta = \frac{1}{2}n\lambda$ (recall that the actual mirror movement is $\frac{1}{2}\delta$)

Modulation Frequency

- *Moving Mirror* moves continuously at a fixed velocity (V_M), so the signal at the detector will oscillate at a related frequency (f):

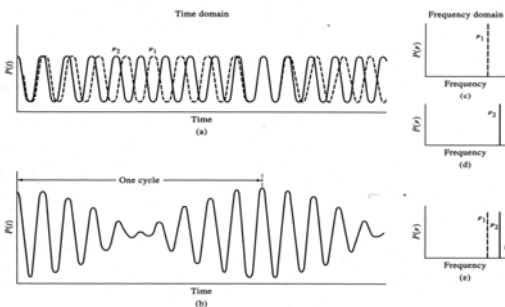
$$f = 2V_M/\lambda$$

Or:

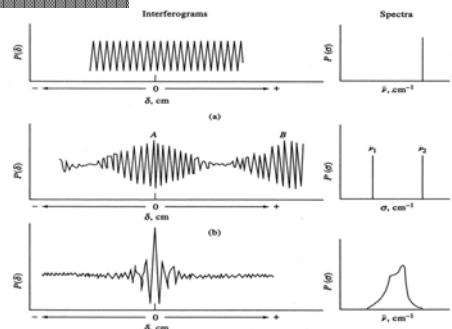
$$f = (2V_M/c)v$$

If $V_M = 0.1 \text{ cm/sec}$, $\lambda = 10 \mu\text{m}$ EMR will be modulated at: $f = 2(1.0 \times 10^{-3} \text{ m/sec})/(10 \times 10^{-6} \text{ m}) = \underline{200 \text{ Hz}}$

To the Frequency Domain!



From Interferogram to Spectrum



Resolution

- It can be shown that for two "barely resolvable" spectral features:

$$\Delta\bar{\nu} = 1/\delta$$

-where δ is the *maximum retardation* attainable with the interferometer.

So, to resolve two spectral features separated by 0.1 cm^{-1} , we need an interferometer with a maximum mirror displacement (x):

$$x = \delta/2 = 10 \text{ cm}/2 = \underline{5 \text{ cm}}$$

37

Why Bother with FT-Interferometry?

1. **Signal-to-Noise Enhancement**
 - **Multiplex Advantage (" Fellgett's Advantage ")**
 - All wavelengths viewed *simultaneously*, so measurement time/resolution element is greater

If measurement is limited by detector noise:

$$\text{S/N enhancement} \propto (n)^{1/2}$$

where n = number of resolution elements

38

Multiplex Advantage: S/N

- So, suppose we acquire a spectrum from $500 - 3500 \text{ cm}^{-1}$ with an effective bandwidth of 1 cm^{-1}

If we spend the same amount of time acquiring the spectrum via FT-Interferometry as we do via scanning spectrometry, we should see a S/N enhancement of:

$n = 3000$ resolution elements

$$(3000)^{1/2} \approx \underline{55\text{-fold}} \text{ S/N enhancement}$$

This is the same as *signal averaging*: like averaging the signal 3000 times at each resolution element

39

Multiplex Advantage: Time

- Suppose we spent 6000 seconds acquiring the spectrum and we really don't need the enhanced S/N:

We can get the same S/N as with a dispersive system in $1/(n)^{1/2}$ of the time

In this case, this means it would take:

$$6000 \text{ sec}/54.8 \approx \underline{110 \text{ sec}}$$

So, 100 minutes (dispersive) versus 2 minutes (FT-interferometry)!

40

Other Advantages

- **Throughput Advantage ("Jacquinot's Advantage")**
 - Increased light throughput (no limiting apertures) *independent of resolution*, results in an increase in S/N.
 - About a 100x S/N enhancement is claimed (not just for systems limited by detector noise)
- 2. **Precise Wavenumber Calibration**
 - due to internal standardization to laser
 - wavenumbers known to about $10^1 - 10^2 \text{ cm}^{-1}$
- 3. **Sheer Elegance and Simplicity!**

41