

CHAPTER SEVEN SOLUTIONS

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Review (a) Newton's Universal Gravitation Law may be written as

$$F = \frac{Gm_1m_2}{r^2} = m_1 \frac{Gm_2}{r^2}.$$

If F is the gravitational force (weight) a mass m_1 located on the surface of a planet experiences, then m_2 is the mass of the planet and r , the distance separating m_1 and m_2 is the radius of the planet R . Comparing the above expression to weight $= m_1g$, we see that the acceleration due to gravity at the surface of the planet is given by

$$g = \frac{Gm_2}{R^2}$$

For Mars, $m_2 = 6.42 \times 10^{23}$ kg, and $R = 3.37 \times 10^6$ m.

Therefore, $g_{\text{Mars}} = (6.673 \times 10^{-11}) \frac{6.42 \times 10^{23}}{(3.37 \times 10^6)^2} = 3.77 \text{ m/s}^2$.

(b) Using $y = v_{0y}t + \frac{1}{2}at^2$ with $v_{0y} = 0$, and $a_y = g_{\text{Mars}}$, we find

$$t = \sqrt{\frac{2y}{g_{\text{Mars}}}} = \sqrt{\frac{2(20.0 \text{ m})}{3.77 \text{ m/s}^2}} = 3.26 \text{ s}.$$

7.3 (a) $\omega = (33 \text{ rev/min})(2\pi \text{ rad/rev})(1 \text{ min}/60 \text{ s}) = 3.5 \text{ rad/s}$.

(b) $\theta = \omega t = (3.5 \text{ rad/s})(1.5 \text{ s}) = 5.2 \text{ rad}$.

7.8 $\omega_f = 2.51 \times 10^4 \text{ rev/min} = 2.63 \times 10^3 \text{ rad/s}$

(a) $\alpha = \frac{\omega_f - \omega_o}{t} = \frac{2.63 \times 10^3 \text{ rad/s} - 0}{3.20 \text{ s}} = 8.22 \times 10^2 \text{ rad/s}^2$.

(b) $\theta = \omega_o t + \frac{1}{2} \alpha t^2 = 0 + \frac{1}{2} (8.22 \times 10^2 \text{ rad/s}^2)(3.20 \text{ s})^2 = 4.21 \times 10^3 \text{ rad}$

7.13 $\theta = \frac{\omega^2 - \omega_o^2}{2\alpha} = \frac{0 - (18.0 \text{ rad/s})^2}{2(-1.90 \text{ rad/s}^2)} = 85.3 \text{ rad}$.

$s = r\theta$ and $r = \frac{1}{2} \text{ diameter} = 1.20 \text{ cm}$,

so $s = (1.20 \text{ cm})(85.3 \text{ rad}) = 102 \text{ cm} = 1.02 \text{ m}$.

7.14 $s = v_o t + \frac{1}{2} at^2 = (17.0 \text{ m/s})(5.00 \text{ s}) + \frac{1}{2} (2.00 \text{ m/s}^2)(5.00 \text{ s})^2 = 110 \text{ m}$.

$s = r\theta$, and $r = 48 \text{ cm} = 0.480 \text{ m}$

Thus, $\theta = \frac{s}{r} = \frac{110 \text{ m}}{0.480 \text{ m}} = 229.2 \text{ rad} = 36.5 \text{ rev}$.

7.21 Let m be the mass of a red corpuscle and let r be the radius of the centrifuge. We are given

$F_c = 4.0 \times 10^{-11} \text{ N} = \frac{mv^2}{r} = mr\omega^2$, so

$$\omega = \sqrt{\frac{F_c}{mr}} = \sqrt{\frac{4.00 \times 10^{-11}}{(3.00 \times 10^{-16})(0.15)}} = 942.8 \text{ rad/s} = 150 \text{ rev/s}.$$

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7.29 (a) At A the forces on the car are the normal force, N , and its weight.

We have $F_c = \frac{mv^2}{r} = N - mg$, or $N = mg + \frac{mv^2}{r}$, which gives

$$N = (500 \text{ kg})(9.80 \text{ m/s}^2) + \frac{(500 \text{ kg})(20.0 \text{ m/s})^2}{10.0 \text{ m}} = 2.49 \times 10^4 \text{ N}.$$

(b) At B we have, $F_c = \frac{mv^2}{r} = mg - N$, or $N = m(g - \frac{v^2}{r})$.

For the vehicle to remain on the track, it is necessary to have

$$N \geq 0 \text{ which means } g \geq \frac{v^2}{r} \text{ or } v = \sqrt{rg}.$$

$$\text{Thus, } v_{\max} = v = \sqrt{rg} = \sqrt{(15.0 \text{ m})(9.80 \text{ m/s}^2)} = 12.1 \text{ m/s}$$

$$\mathbf{7.39} \text{ (a) } PE = -\frac{GM_{\text{Em}}}{r} = \frac{(6.67 \times 10^{-11} \text{ Nm}^2/\text{kg}^2)(5.98 \times 10^{24} \text{ kg})(100 \text{ kg})}{(8.38 \times 10^6 \text{ m})}$$

$$= -4.76 \times 10^9 \text{ J}$$

$$\text{(b) } F = \frac{GM_{\text{Em}}}{r^2} = \frac{(6.67 \times 10^{-11} \text{ Nm}^2/\text{kg}^2)(5.98 \times 10^{24} \text{ kg})(100 \text{ kg})}{(8.38 \times 10^6 \text{ m})^2}$$

$$= 5.68 \times 10^2 \text{ N}$$

7.50 (a) At the bottom of the swing, $T - mg = \frac{mv^2}{L}$, or

$$T = mg + \frac{mv^2}{L} = (0.400)(9.80) + \frac{(0.400)(3.00)^2}{0.800} = 8.42 \text{ N}.$$

(b) From conservation of energy, we may write

$(PE_g)_{\text{at top of swing}} = (KE)_{\text{at bottom}}$, or

$$mgL(1 - \cos\theta_{\max}) = \frac{1}{2} mv^2_{\text{bottom}}, \text{ giving}$$

$$\cos\theta_{\max} = 1 - \frac{v^2_{\text{bottom}}}{2gL} = 1 - \frac{(3.00 \text{ m/s})^2}{2(9.80 \text{ m/s}^2)(0.800 \text{ m})} = 0.426.$$

Thus, $\theta_{\max} = 64.8^\circ$.

(c) At $\theta = \theta_{\max}$, the pendulum is at rest, so the radial force

$$F_c = \frac{mv^2}{L} = T - Mg\cos\theta_{\max} = 0.$$

$$\text{Thus, } T = Mg\cos\theta_{\max} = (0.400 \text{ kg})(9.80 \text{ m/s}^2)\cos 64.8^\circ = 1.67 \text{ N}.$$

7.57 Using Kepler's Third law $T^2 = ka^3$, we have

$$(75.6)^2 = \left(\frac{0.57 + x}{2}\right)^3$$

The farthest distance $x = 2(75.6)^{2/3} - 0.57 = 35.2 \text{ A.U.}$ (near the orbit of Pluto)

ANSWERS TO EVEN ASSIGNED CONCEPTUAL QUESTIONS

4. The speedometer will be inaccurate. The speedometer measures the number of revolutions per second of the tires.